

SAMPLE SOLUTIONS EXERCISE 8

EXERCISE 8.1: ENTROPY OF MANY HARMONIC OSCILLATORS (6P)

In this exercise we want to determine the entropy of N classical three-dimensional harmonic oscillators in the microcanonical ensemble as a function of the energy E .

(a) Compute the phase space volume

$$V(E, N) = \int_{\mathcal{H} \leq E} \prod_{i=1}^{3N} dq_i dp_i \quad \text{where} \quad \mathcal{H} = \sum_{i=1}^{3N} \left(\frac{p_i^2}{2m} + \frac{m\omega^2 q_i^2}{2} \right).$$

Hint: Relate this integral to the volume of a $6N$ -dimensional sphere. (2P)

(b) Determine the phase space volume in an infinitesimal energy shell $[E, E + \delta E]$, given by (1P)

$$\delta V(E, N, \delta E) = \delta E \frac{\partial V(E, n)}{\partial E}.$$

(c) As a rough approximation, let us assume that each (quantum) state occupies a volume $(2\pi\hbar)^{3N}$ in the shell. Use Stirling's formula to approximate the entropy of the states in the energy shell $H(E, N, \delta E) = k_B \ln |\Omega(E, N, \delta E)|$, keeping only the terms which grow at least linearly with N . (2P)

(d) Calculate the temperature of the system which is defined as $T^{-1} = \beta = \frac{\partial H}{\partial E}$. (1P)

SAMPLE SOLUTION

(a) We introduce $6N$ dimensionless coordinates by

$$x_i = \begin{cases} \frac{1}{\sqrt{2m}} p_i & \text{for } i = 1, \dots, 3N \\ \sqrt{\frac{m\omega^2}{2}} q_{i-3N} & \text{for } i = 3N + 1, \dots, 6N \end{cases}$$

turning the Hamilton function into

$$H = \sum_{i=1}^{3N} (x_i^2 + x_{i+3N}^2) = \sum_{i=1}^{6N} x_i^2.$$

Since $dx_i dx_{i+3N} = \frac{\omega}{2} dq_i dp_i$ the volume turns into

$$V(E, N) = \left(\frac{2}{\omega} \right)^{3N} \int_{H \leq E} \prod_{i=1}^{6N} dx_i = \left(\frac{2}{\omega} \right)^{3N} \frac{\pi^{3N} E^{3N}}{\Gamma(3N + 1)} = \frac{(2\pi E)^{3N}}{\omega^{3N} (3N)!},$$

where we used the formula for the volume of an n -dimensional sphere (see e.g. Wikipedia) which here has the radius $r = \sqrt{H} \leq \sqrt{E}$.

(b) We simply take the derivative with respect to E :

$$\delta V(E, N, \delta E) = \delta E \frac{\partial V(E, n)}{\partial E} = \frac{3N\delta E}{E} \frac{(2\pi E)^{3N}}{\omega^{3N} (3N)!}$$

(c) The entropy in the energy shell is estimated by

$$H = k_B \ln |\Omega| = k_B \ln \left(\frac{\delta V(E, N, \delta E)}{(2\pi\hbar)^{3N}} \right) = k_B \ln \left(\frac{3N\delta E}{E} \frac{E^{3N}}{(\hbar\omega)^{3N} (3N)!} \right)$$

This can be rewritten as

$$H = k_B \left[\ln \left(\frac{3\delta E}{E} \right) + \ln N + 3N \ln \left(\frac{E}{\hbar\omega} \right) - \ln((3N)!) \right]$$

We now apply Stirling's formula $n! \approx \sqrt{2\pi n}(n/e)^n$

$$H = k_B \left[\ln \left(\frac{3\delta E}{E} \right) + \ln N + 3N \ln \left(\frac{E}{\hbar\omega} \right) - \frac{1}{2} \ln(6\pi N) - 3N \ln(3N/e) \right].$$

We sort the terms by their significance:

$$H = k_B \left[3N \left(\ln \left(\frac{E}{3N\hbar\omega} \right) + 1 \right) + \frac{1}{2} \ln N + \ln \left(\frac{3\delta E}{E} \right) - \frac{1}{2} \ln(6\pi) \right].$$

In the large N limit, the last three terms can be neglected, leading to the approximation

$$H \approx 3Nk_B \left[\ln \left(\frac{E}{3N\hbar\omega} \right) + 1 \right],$$

where the '+1' comes from the 'e' in the last term.

(d) Compute the partial derivative:

$$T^{-1} = \beta = \left(\frac{\partial H}{\partial E} \right)_N = \frac{3Nk_B}{E}.$$

This confirms that the average energy per 3D oscillator is indeed $3k_B T$.

EXERCISE 8.2: FREE ENERGY OF A PERTURBED SYSTEM

(6P)

A classical system in thermal equilibrium with a heat bath at temperature T is described by an energy function $E^{(0)} : \Omega^{\text{sys}} \rightarrow \mathbb{R} : s \rightarrow E_s^{(0)}$. Let $Z^{(0)} = \sum_s e^{-\beta E_s^{(0)}}$ be the partition sum of the system.

- Assume that the energy function is perturbed by $E_s^{(0)} \rightarrow E_s = E_s^{(0)} + \lambda E_s^{(1)}$ with $\lambda \ll 1$. Compute the corresponding partition sum $Z(\lambda)$ as a power series in λ . (2P)
- What is the mathematical meaning of $Z(\lambda)$? Hint: Try to compute $\frac{d^n}{d\lambda^n} Z(\lambda) \Big|_{\lambda=0}$ (1P)
- Prove the general statement that in the canonical ensemble the free energy $F = E - TH$ (more precisely: $\langle F \rangle = \langle E \rangle - T \langle H \rangle$) is given by $F = -T \ln Z$. (1P)
- Apply (c) to (a),(b) in order to compute $F(\lambda)$ as a power series in λ up to second order. What is the mathematical meaning of $F(\lambda)$? (2P)

SAMPLE SOLUTION

- (a) In the canonical ensemble the partition sum reads $Z^{(0)} = \sum_s e^{-\beta E_s^{(0)}}$, where $\beta = 1/T$ is the inverse temperature and the sum runs over all system configurations $s \in \Omega^{\text{sys}}$. In the unperturbed case, the expectation value of an arbitrary random variable X is given by

$$\langle x \rangle_0 = \frac{1}{Z^{(0)}} \sum_s e^{-\beta E_s^{(0)}} x_s.$$

Setting $x_s = e^{-\beta \lambda E_s^{(1)}} = 1 - \beta \lambda E_s^{(1)} + \mathcal{O}(\lambda^2)$ we simply get

$$\begin{aligned} Z(\lambda) &= \sum_s e^{-\beta E_s} = \sum_s e^{-\beta(E_s^{(0)} + \lambda E_s^{(1)})} = Z^{(0)} \langle e^{-\beta \lambda E^{(1)}} \rangle_0 \\ \Rightarrow Z(\lambda) &= Z^{(0)} \left\langle \sum_{n=0}^{\infty} \frac{(-\beta \lambda E^{(1)})^n}{n!} \right\rangle_0 = Z^{(0)} \sum_{n=0}^{\infty} \frac{\lambda^n}{n!} \langle (-\beta E^{(1)})^n \rangle_0. \end{aligned}$$

- (b) We have

$$Z(\lambda) = Z^{(0)} \sum_{n=0}^{\infty} \frac{\lambda^n}{n!} m_n = Z^{(0)} M(\lambda)$$

where $m_n = \langle (-\beta E^{(1)})^n \rangle_0$, thus $Z(\lambda)$ may be interpreted as the moment-generating function $M(\lambda)$ of the perturbation $-\beta E^{(1)}$ multiplied by $Z^{(0)}$.

- (c) We start with the definition of the Shannon entropy

$$\begin{aligned} H &= - \sum_s P_s \ln P_s = - \frac{1}{Z} \sum_s e^{-\beta E_s} \ln \frac{e^{-\beta E_s}}{Z} \\ &= \frac{1}{Z} Z \ln Z + \frac{\beta}{Z} \sum_s e^{-\beta E_s} E_s = \ln Z + \beta \langle E \rangle. \end{aligned}$$

This implies

$$\ln Z = H - \beta E = -\beta(E - TH) = -\beta F \quad \Rightarrow \quad F = -T \ln Z$$

- (d) Writing down the free energy we realize that

$$F(\lambda) = -T \ln Z(\lambda) = -T \ln Z^{(0)} - T \ln \sum_{n=0}^{\infty} \frac{\lambda^n}{n!} m_n = F^{(0)} - T K(\lambda),$$

where $K(\lambda)$ is the cumulant-generating function of the perturbation $-\beta E^{(1)}$. The zeroth cumulant is always zero, the first one is the mean, and the second cumulant is the variance. Thus we can conclude that

$$\begin{aligned} F(\lambda) &= F^{(0)} - \frac{1}{\beta} \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \kappa_n \\ \Rightarrow F(\lambda) &= F^{(0)} + \lambda \langle E^{(1)} \rangle_0 - \frac{\beta \lambda^2}{2} \langle (E^{(1)} - \langle E^{(1)} \rangle)^2 \rangle + \mathcal{O}(\lambda^3). \end{aligned}$$

So the free energy is basically the cumulant-generating function of $-\beta E^{(1)}$.

($\Sigma = 12\text{P}$)